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# Modal noise mitigation in few-mode fibers



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#### CONTEXT

- NIRPS (Near InfraRed Planet Searcher) is an <u>AO-assisted fiber-fed RV NIR spectrograph</u> for La Silla
- YJH-band RV spectroscopy focusing on rocky planets around M-stars
  - 1 m/s radial velocity precision
  - Planetary atmospheres characterization

•	Mode	Fiber at telescope	Fiber at spectrograph	R
	AO with 0.4" fiber	OCT 29 $\mu$ m	OCT 29 $\mu$ m	100.000



CaF2/S-FTM16

Seeing-limited with 0.9" fiber

OCT 66 $\mu$ m RECT 33x132 $\mu$ m

- Simultaneous observations with HARPS
- First light end of 2019

#### **COUPLING EFFICIENCY**

Blind+, 2017 (AO4ELT)

75.000

Thanks to the AO system, star light is coupled to an 0.4" fiber with high coupling efficiency

- Coupling efficiency can outperform seeing-limited instrument by 30 to 50% (Fig. 1)
- Gain of 1-2 magnitude over single-mode solution



# FIBER SCRAMBLING WITH DM

- DM is used to couple light to the various modes of the fiber
- Mode-selective coupling is possible to some extent, but is not photon efficient



Fig. 5 – Fiber filling factor when applying a Zernike OPD on DM. Grey area shows 10% coupling loss from optimal.

#### **MODAL NOISE IN FEW-MODE FIBERS**

- 0.4" fiber in YJH bands means 10 to 50 modes are guided (vs 1000s for HARPS)
- NIRPS fibers work in the few-mode regime
  - Ray tracing does not apply anymore!!! (Fig. 2)
- <u>Modal noise</u>: variation of fiber output illumination, caused by changing input illumination condition (e.g. tracking error) or external factors (e.g. temperature)
- Modes well separated: low scrambling efficiency and high modal noise , ~20 times higher than for HARPS (Fig. 3)



Fig. 2 - Near-field of 50-mode ( $\lambda$ =950nm) and 15-mode ( $\lambda$ =1650nm) octagonal fiber when injecting in core center.



Fig. 3 – Photocenter instability in the near-field of the 0.4" fiber

# Tip-tilt scan is the most efficient term, with a scrambling gain of ~10

hp-tilt scan is the most efficient term, with a scramping gain of "TO

### **FIBER STRETCHER**

- Fiber Stretcher modulates the phase between modes → uniformizes interference pattern (Fig. 8)
- 7mm stretch  $\Leftrightarrow \Delta T=25K$  over 50m of fiber
- Makes illumination independent of external factors such as temperature, varying stress on fiber during tracking, etc.

#### Performance

- Throughput > 90%
- Scrambling gain ~ 4





Fig. 6 - Prototype stretcher with 7mm total modulation amplitude @20Hz



#### **DOUBLE SCRAMBLER**

- Near-field and far-field are not independent anymore —
- Few-mode fiber intrinsic scrambling efficiency is low
- Near-field to far-field modes mismatch costs 10-20% in throughput
- Performance estimated by injecting far-field measurements into spectrograph optical design (Fig. 4)

The double scrambler does not improve pupil stability in the few-mode regime

pupil stability in the few-mode regime



→ double scrambler not effective

Fig. 4 – Standard feedthrough and double scrambler sensitivity to different sources of error

Fig. 7 – Modal noise in m/s VS stretch amplitude.



Fig. 8 – Near-field output image (top) and radial profiles vs injection position (bottom).

## **RADIAL VELOCITY BUDGET**

	Scrambling gain	RV
Octagonal fiber 0.4" [980-1800nm]	300	< 10 m/s
Double scrambler	1	< 10 m/s
DM Tip-Tilt	10	< 1.0 m/s
Stretcher	4	< 25 cm/s
NIRPS total		< 25 cm/s